## Comparative Planetology

The study of the solar system by examining and understanding the similarities and differences among worlds

We'll start with the terrestrial planets

## Comparative Planetology

# 1.Interiors <br> 2.Geologic activity \& processes 

## 3. Surfaces

4.Atmospheres 5. Magnetic fields

These are the things we need to look at to understand how (terrestrial) planets form. We'll start with the Earth, and compare from there.

## Comparative Planetology

## Investigation:

Using the books \& internet, come up with some basic characteristics
of "your" planet


We'll start with the terrestrial planets

## Comparative Planetology

$$
\begin{aligned}
& \text { Before we get too far: } \\
& \text { Most of the water (etc) in the } \\
& \text { inner solar system arrived during } \\
& \text { heavy bombardment }
\end{aligned}
$$

The amount of water and other light molecules/elements that we have here on Earth, Venus and Mars all came from comet impacts during the heavy bombardment. A small amount probably came from outgassing as it may have been trapped inside some rocks (we do see ice on the asteroids: remember if the dust is thick enough we can "shield" some ice so that it sticks around in the inner solar system, just not a lot).

## Internal Structure I

## Differentiation:

## Sorting of planetary material by density.



Liquids differentiate by density as well. Oil floats on water.
The interiors of the terrestrial planets are hot. accretion, released gravitational potential, radioactive decay.
Rocks are viscous (meaning that they can flow).
Note that large bodies are spherical (lots of gravity eventually causes the rock to flow into a spherical shape)
Smaller bodies (asteroids) are misshapen because there isn't enough gravity to deform them.
Hot rocks flow better than cold rocks
Heavy stuff tends to sink to the center of the planet (ie iron and heavy metals)
Lighter stuff (silicates, rocks) rises to the top.

## Internal Structure I



## Earth and Moon



Here is an example of the kinds of things we will be doing: looking at both the Earth and the moon to figure out similarities and differences.


The lithosphere contains the crust and part of the mantle.
Higher temperatures deeper in the planet allow the rock there to be more fluid and the lithosphere essentially 'floats' on top of the mantle.
Large planets have more internal heat and thus a thinner lithosphere.
Small planets lose their heat more quickly and have a thicker lithosphere.

Mercury has a GIANT core. Was it made from more metal rich planetesimals? The leading hypothesis is that it suffered some large impact in the past that blasted away part of its mantle. (So where did it go??)

Earth and Venus have very similar interiors.

Mars has a very thick lithosphere, because it is small. It may have a very small amount of semi-plastic mantle, however the lithosphere is so thick that there is no current tectonic activity. Its core is also fairly frozen: we can tell because there is not a global magnetic field any more (though there used to be one).

The moon is all lithosphere and frozen core. Notice that it has a higher
lithosphere to core ratio than Earth/Venus/Mars, which is probably due to the big impact that smacked off part of the Earth, which then coalesced into the moon. (the impact would have knocked off more lithosphere/mantle than core material: what core is there may be part of the original impactor, though it's hard to say for certain)

NOTE: This differentiation is not the same as differentiation in the solar nebula!

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## Magnetic Fields



You can see the shape of the magnetosphere in this picture. It's asymmetric because of the solar wind hitting it. Inside our magnetosphere we are shielded from lots of high energy protons from the solar wind. This also helps satellites, except those that go far enough out that they are outside the magnetopause.

## Magnetic Fields

## Spinning molten iron core



Free electrons in the molten core are dragged around by the rotating earth giving rise to magnetic fields.
Of the terrestrial planets, only the Earth has a strong magnetic field.
It protects us from harmful cosmic rays and diverts the solar wind around us. Also partially responsible for aurora
In mars and mercury the core is likely solid or mostly solid. Venus rotates too slowly to have a magnetic field, even though its core is most likely molten.

## Aurora



## Magnetic Fields



Mercury: what's up with that? We don't know; we only just found out about it!

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## Geology



What kinds of geologic activity can we think of?

## Earth

## Lots of evidence for ongoing geologic activity



## Geologic activity

## Types:

- Earthquakes
- Volcanoes
- Continental Shift
- Sea floor spreading
- Mountain Formation
- Erosion


## Causes:

- Plates shifting
- High Temperatures, pressure, magma surfacing
- Plates floating
- Sea floor spreading
- Water, wind, Glaciers, people, lava

Not all of these happen on all planets!

## What causes geologic activity?

## Heating

- Accretion
- Differentiation
- Radioactive decay

Cooling

- Convection
- Conduction
- Radiation

The amount of heat a planet currently has drives its geologic activity. These all (roughly) relate to the planet's SIZE.

## Tectonics

Tectonics: The action of internal forces on the lithosphere to create surface features.


Earth and Venus have ongoing tectonic activity.
Mars had such activity in the past

## Plate Tectonics

## Great plates of crust slowly move around the Earth



## Lecture Tutorial (p. 101)

Subduction is Where one plate dives under another.
Where plates are moving apart rifts form. Lava often wells up to fill in the rift.
Mountains form where the plates meet.
Earthquakes happen when plates stick and snap loose suddenly

## Tectonics

## Are plate tectonics the only type of tectonic activityidiNOPE!



Stagnant Lid Tectonics
Planetary contraction

NOPE! Earth is the only planet with plate tectonics but not the only planet with tectonics!
**Why is Earth the only planet with plate tectonics?** (question to be answered later, but pose it here)

## Mercury: Contraction



Some impact craters do show lava in-fill (much like the lunar maria), and as that cooled, we see cracks. The lava plains are the largest volcanic features on Mercury and likely have thicknesses of several kilometers thick. Mercury doesn't currently have any tectonic activity.

## Mercury

## Thrust faults due to squeezing



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## Tectonic Activity



# Olympus Mons 

700 km diameter

25 km high

Volcanoes on Mars can grow very tall because of the low surface gravity... Olympus Mons is the tallest, and is more than 2.5 times higher than Mount Everest ( 8.8 km high)! There are three other large volcanoes in the Tharsis region. There are lots of other volcanoes too, but these four are the largest. Because of the huge weight on the crust, we have...

## Mars: Extension

## Valles Marineris

## Much grander than the Grand Canyon

4000 Kilometers long (the continental US is about $3,000 \mathrm{~km}$ long)
50 to 100 Kilometers wide
10 km deep ( 7 times deeper than the grand canyon)

Likely formed by tectonic stresses - due to Olympus Mons and three other large volcanoes to the west of the crack. Also note the giant outflow region to the north side... probably a catastrophic outflow caused this.

There used to be volcanic activity in the distant past... Mars is cool now

## Venus

## Similar to Earth in size



## Shrouded in clouds

These images are created with radar.
Radio waves penetrate the clouds. Time the echo of the emitted radio pulse to determine the topology of the landscape.


But, we don't actually have plate tectonics. Why not? Because they're no water - which makes it much harder for rocks to "flow" like they do here on Earth. So what do we have going on instead? As the lava covers the surface, it gets heavy and sinks down into the mantle. Other areas upwell... so plains are sinking and areas of uplift are rising. This causes long linear features (fossae) due to extension and compression as well.

## Tectonic Effects: Volcanism



> The type of volcano depends on the viscosity of the lava

The interior of the planet is not all molten lava. The lava is in hot pockets in the lithosphere.
Molten lava can erupt to the surface significantly changing the surface of a planet by erasing older features.
The planetary interior must be hot to have active volcanism. **How do we heat the interior?** (one minute write/think pair share)

## Volcanism

## Shield volcanoes have moderately viscous flows and tend to spread out



Mauna Loa


Olympus Mons

## Stratovolcanoes

## Highly viscous lava flows build very tall volcanic mountains



Mt. Fuji

Volcanic activity requires a hot planetary interior

## Venus




Shield Volcanoes

Low cratering.
It looks like the surface was completely recreated approximately a billion years ago.
Evidence of recent tectonic activity and volcanic activity, by the sulfuric acid in the atmosphere - if there were no current volcanoes, it would become locked up in the surface rocks.
There does not appear to be any erosion despite the thick atmosphere

## The Moon



## Signs of past geologic activity: Maria

Chemical composition much like the Earth's crust

Heavily cratered during heavy bombardment period.
Radioactive decay slowly heated the interior until it was hot enough for some volcanism
Volcanic eruptions created the maria
The dark color of the maria is due to iron rich rock from the interior.
The interior re-solidified.
Surface continued to crater
No atmosphere, so no erosion. No tectonic activity
The surface is powdery... impacts from micro-meteorites.

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## Surfaces: Cratering



## Impactor strikes the surface at a high velocity

Surface rock is vaporized as the kinetic energy is released

Material is pushed aside creating the rim

Typical impact velocities are between 10,000 and 100,000 km/hour Craters are much wider than the impactor...
Material is thrown a large distance in a spray around the crater.

## Cratering

## Craters reveal historical information



Small craters outnumber large craters.
Large craters are overlapped by smaller ones (the ones on top are newer) There are few impacts in the Moon's Maria. (so they are newer)

Craters provide a way to figure out what features are older and what features are newer.

## Mercury



Heavily cratered
signs of tectonic activity in the distant past, smooth lava plains like the moons maria
Evidence for a very large solid iron core.
It is now geologically dead (no tectonic activity)

At the poles are very likely water ice deposits.

## Mercury



The crater is about 1550 km across and the ringing mountains are about 2 km high. The floor is filled with lava plains (like the Moon's maria). As with many big impacts there are concentric rings around it. There are some other weird things inside the crater, like what look like extension faults. MESSENGER should be able to sort out what some of these things are.


All of these planets have fewer craters than Mercury and the Moon. Why? Hint: Venus has a thick atmosphere, Earth is medium, and Mars has a very thin atmosphere.

## Impacts and Mass Extinctions

## Meteor Crater

Caused by a 50-meter
asteroid about
50,000 years ago


65 million years ago,
many species disappeared.

Sedimentary rock layer from that time shows a number of features from that event

We know large objects have hit the Earth.
Iridium, Osmium, Platinum, grains of "shocked ctuarez", spherical rock droplets, soot from forest fires
In 1991, a 65 million year old impact crater was found on the coast of Mexico.
200 km in diameter
implies an asteroid size of about 10 km across
called the Chicxulub crater
Coincides with the mass extinction of the dinosaurs
An impact of this size would have caused:
debris in atmosphere blocks sunlight; plant die...animals starve poisonous gases form in atmosphere

## Erosion

Erosion: Wearing away and/or transportation of rock by wind, water, or ice


On a planet with a significant atmosphere, erosion is an ongoing process. **Which planets have a "significant" atmosphere?**

## Mars



Nearly 24 hour day, tilt very similar to the Earth
Light cratering
Half the radius of the Earth
Plenty of evidence of past geologic processes but no active volcanism (other, smaller, processes are happening though)

## Evidence for Water



Hematite Blueberries


## Many apparent flows

Blueberries and other salt deposits suggest that the surface was once bathed in water.
Many surface features are strongly suggestive of liquid flows - we definitely see flows currently happening but whether or not it's liquid or slush/slurry is unclear
**Where is the water now?**
At this point we're certain there was once liquid water on Mars - this finding is extremely recent!!

# Extra Terrestrial Landscapes 



Go to the mars rovers page and see LOTS of pretty pictures and read lots of stuff about mars.

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